University of Alabama in Huntsville

Research Institute

6-1-1966

Unsteady Heat Transfer from a Rotating Disk Due to a Step Change in Local Wall Heat Flux with Large Suction UARI Research Report No. 27

D. R. Jeng

Y. V. Subba Rao

Follow this and additional works at: https://louis.uah.edu/research-reports

Recommended Citation

Jeng, D. R. and Subba Rao, Y. V., "Unsteady Heat Transfer from a Rotating Disk Due to a Step Change in Local Wall Heat Flux with Large Suction UARI Research Report No. 27" (1966). *Research Institute*. 20. https://louis.uah.edu/research-reports/20

This Report is brought to you for free and open access by LOUIS. It has been accepted for inclusion in Research Institute by an authorized administrator of LOUIS.

Casters

UARI Research Report No. 27

UNSTEADY HEAT TRANSFER FROM A ROTATING DISK DUE TO A STEP CHANGE IN LOCAL WALL HEAT FLUX WITH LARGE SUCTION

by

D. R. Jeng and Y. V. Subba Rao

The research for this report was supported by the National Aeronautics and Space Administration under research grant NsG-381

UNIVERSITY OF ALABAMA RESEARCH INSTITUTE Huntsville, Alabama June 1966

October 14, 1965

Masters

Office of Grants & Contracts Gode SC National Aeronautics & Space Administration Washington 25, D.C.

Attention: Dr. T. L. K. Smill

Dear Dr. Smill:

Enclosed are five (5) copies of the University of Alabama Research Institute Research Report No. 27 entitled, "Unsteady Hest Transfer From a Rotating Disk Due to a Stop Change in Local Wall Heat Flux with Large Suction", by D. R. Jeng and Y. V. Subba Rao. The research for this report was supported by the National Aeronautics and Space Administration Grant MsG-381.

Sincerely yours,

NP.N.

W.P. Watts Administrative Manager

Enclosures (5)

co: - Director for Candractor & Crasts, University of Alabama (1) Director, Research Division, National Acronautics & Space Administration October 12, 1966

Commanding General U. S. Army Missile Command Redstone Arsenal, Alabama 35809

Attention: AMSMI-G/Mr. R. R. Finney

References: (a) Final Technical Report on Army Missile Command Contract DA-01-021-AMC-14042(Z) entitled "Multi-Rail Launcher with Six Degrees of Freedom"

(b) AMXBI-PIO May 25, 1965

Dear Sir:

eat

and

Approval for open distribution of reference (a) as a UARI Research Report is requested.

Pursuant to the requirements of reference (b), we are forwarding two copies of the report.

Sincerely,

Inan

W. P. Watts Administrative Manager

(ETR)

cc: Chief of Information, Department of the Army Attn: Civil Liaison Division

UARI Research Report No. 27

UNSTEADY HEAT TRANSFER FROM A ROTATING DISK DUE TO A STEP CHANGE IN LOCAL WALL HEAT FLUX WITH LARGE SUCTION

by

D. R. Jeng and Y. V. Subba Rao

The research for this report was supported by the National Aeronautics and Space Administration under research grant NsG-381

UNIVERSITY OF ALABAMA RESEARCH INSTITUTE Huntsville, Alabama June 1966

TABLE OF CONTENTS

| Abstract | iii |
|---|-------|
| Nomenclature | iv,v |
| 1.0 Introduction | 1 |
| 2.0 Governing Equations for Incompressible Flow | 3 |
| 3.0 Solution Method | 5 |
| 4.0 Numerical Results and Discussion | 24 |
| 5.0 A Special Case of Compressible Flow | 25 |
| References | 26,27 |
| Table 1 Values of the Dimensionless Temperature, $\theta_{ss}(0)$, for Large Su | ction |
| Table 2 Values of S(0) Values of Q(0) | |
| Table 3 Roots of the Cubic Equation | |
| Fig. 1 Coordinate and Velocity Nomenclature | |
| Fig. 2 Dimensionless Temperature Response at the Surface of a Rotatin Disk Due to a Step Change in Local Heat Flux | g |
| Fig. 3 Dimensionless Temperature Response at the Surface of a Rotatin Disk Due to a Step Change in Local Heat Flux | g |
| Fig. 4 Dimensionless Temperature Response at the Surface of a Rotatin Disk Due to a Step Change in Local Heat Flux | g |
| Fig. 5 Dimensionless Temperature Response at the Surface of a Rotatin Disk Due to a Step Change in Local Heat Flux | g |
| Fig. 6 Dimensionless Temperature Response at the Surface of a Rotatin Disk Due to a Step Change in Local Heat Flux | g |
| Fig. 7 Dimensionless Temperature Response at the Surface of a Rotatin Disk Due to a Step Change in Local Heat Flux | g |
| Fig. 8 Similarity Presentation of the Results Shown in Figures 2 to 7 for $P_r = 0.1$ to 10 and $a_0 = 1.5$ to 4.0 | |

Abstract

An analysis is made for the unsteady heat-transfer, due to a time dependent wall heat flux, from an infinite disk, rotating in still fluid with large suction. The procedure begins with a consideration of the thermal response caused by a step change in the local wall heat flux. It is found that the response time varies inversely with the Prandtl number, the angular velocity and the second power of the suction parameter.

Incompressible flow with dissipation is considered first. For an ideal gas with constant Prandtl number whose viscosity varies linearly with temperature, the incompressible results may be, if dissipative effect can be ignored, directly applied provided that all properties in the solution are replaced by their wall values.

| Nomenclature | |
|----------------|---|
| a _o | Mass transfer parameter, $-\frac{w_w}{(v w)^{1/2}}$, dimensionless |
| с _р | Specific heat at constant pressure, J/kg ^O K |
| erf x | error function, $\frac{2}{\sqrt{\pi}} \int_{0}^{\infty} e^{-5^2 d5}$ |
| F | Anners, here? |
| с } н | function defined by (6) |
| erfc x | 1 - erf x |
| k | thermal conductivity, J/m sec ^O K |
| р | parameter in Laplace transform |
| р | static pressure, n/m ² |
| Pr | Prandtl number, dimensionless |
| q _w | Wall surface heat flux, J/sec m ² |
| Q(IJ) | function defined in (11) or (15) |
| r R | radial coordinate for rotating disk, m dimensionless radial coordinate, $r(\frac{\omega}{v})^{1/2}$ |
| s(ŋ) | function defined in (11) or (15) |
| t | time, sec. |
| T | temperature, ^o K |
| u | velocity parallel to the surface of a disk, m/sec |
| v | velocity in ϕ direction, m/sec |
| w | velocity normal to the surface of a disk, m/sec |

iv

| Z | coordinate measuring distance normal to a disk, m |
|------|--|
| η | dimensionless coordinate $\eta = Z(\frac{\omega}{V})^{1/2}$ |
| θ | dimensionless temperature defined by (11) or (15) |
| μ | viscosity, kg/m sec |
| ν | kinematic viscosity, m ² /sec |
| ρ | density, kg/m ³ |
| τ | dimensionless time, $\frac{\omega t}{P_r}$ |
| ω | angular velocity, rad/sec |
| φ | the angular position |
| 1(t) | Heaviside unit operator; = 0 for $t < 0$, and = 1 for $t \ge 0$ |
| | |

0

Subscripts

| i | initial conditions | | | | | |
|----|---------------------------|--|--|--|--|--|
| SS | steady state | | | | | |
| w | condition at wall surface | | | | | |
| 00 | free stream condition | | | | | |

and and astimization belocity joined to cover a wide range of Frankt, ambers to the sublime velocity becomes sufficiently large, a closed-form there an in intrined. The analysis for this case was sloo made in it has for a step charge in wall compensators. In the present report, mentions the same conjuncty heat transfer problem but with the stop ings: in wall heat flow. The withod of solution adapted in the present alysis also be presented to the used in [9] but the scalysis become me proportioned. We firsh consider a case of incompressible flow. A

1.0 Introduction

Unsteady heat transfer due to a time prescribed wall temperature or heat flux has been a subject of interest for many years. Sparrow and Gregg [1,2] investigated the laminar forced convection heat transfer from a compressible fluid to a flat plate with uniform, but time dependent, surface temperature. Cess [3] and Riley [4] examined the same problem for incompressible flow, Goodman [5] and Adams and Gebhart [6] have employed the heat balance integral technique to obtain approximate solutions for the flat plate problem. Sparrow [7] reported an approximate analysis for the unsteady, two dimensional stagnation point heat transfer. He employed the integral technique for which a third degree polynomial was chosen for the unsteady temperature profile. Subsequently Chao and Jeng [8] published an analysis for the unsteady heat transfer at a two-dimensional and axisymmetrical front stagnation point due to an arbitrarily prescribed wall temperature or heat flux. The analysis was extended by Jeng [9] to a three-dimensional magnetohydrodynamic stagnation point flow with simultaneous suction or blowing. The heat transfer from rotating bodies is of technological interest and has attracted the attention of several researchers. The steady heat transfer from a rotating disk was first studied by Millsaps and Pohlhausen [10] and later by Sparrow and Gregg [11] who also investigated the effect of blowing and suction [12]. Cess and Sparrow [13] were probably the first to analyze the unsteady heat transfer from a rotating disk. The problem was later re-examined by Jeng [9] who also included the effects of mass transfer. Two asymptotic solutions, respectively valid for small and large times, are found and satisfactorily joined to cover a wide range of Prandtl numbers. When the suction velocity becomes sufficiently large, a closed-form solution can be obtained. The analysis for this case was also made in [9] but for a step change in wall temperature. In the present report, we investigate the same unsteady heat transfer problem but with the step change in wall heat flux. The method of solution adopted in the present analysis closely parallels to that used in [9] but the analysis becomes more complicated. We first consider a case of incompressible flow. A solution describing the entire time history of the non-steady temperature

1

field has been obtained. For an ideal gas with constant Prandtl number whose viscosity varies linearly with temperature, it is shown that the incompressible result may be directly applicable to compressible case.

2.0 Governing Equations for Incompressible Flow

Consider an infinite disk rotating in an infinite mass of fluid about an axis normal to its own plane and at a constant angular velocity W. Fig. 1 illustrates the cylindrical coordinate (r, ϕ, z) and the corresponding velocity components (u, v, w) appropriate for the problem. From physical considerations, one sees that the velocity and temperature field would be independent of ϕ , if the thermal condition at the disk surface is also independent of ϕ . Under the assumption of steady, incompressible flow with constant properties, the governing equations are (with dissipation included):

Continuity:

$$\frac{\partial u}{\partial r} + \frac{u}{r} + \frac{\partial w}{\partial z} = 0$$
(1)

Momentum:

$$u \frac{\partial u}{\partial r} - \frac{v^2}{r} + w \frac{\partial u}{\partial z} = -\frac{1}{\rho} \frac{\partial p}{\partial r} + v \left[\frac{\partial^2 u}{\partial r^2} + \frac{1}{\partial r} \left(\frac{u}{r} \right) + \frac{\partial^2 u}{\partial z^2} \right]$$
(2.a)

$$u \frac{\partial w}{\partial r} + w \frac{\partial w}{\partial z} = -\frac{1}{\rho} \frac{\partial p}{\partial z} + \nu \left[\frac{\partial^2 w}{\partial r^2} + \frac{1}{r} \frac{\partial w}{\partial r} + \frac{\partial^2 w}{\partial z^2} \right]$$
(2.c)

Energy:

$$\frac{\partial \mathbf{T}}{\partial \mathbf{t}} + \mathbf{u} \frac{\partial \mathbf{T}}{\partial \mathbf{r}} + \mathbf{w} \frac{\partial \mathbf{T}}{\partial \mathbf{z}} = \alpha \left[\frac{1}{\mathbf{r}} \frac{\partial}{\partial \mathbf{r}} \left(\mathbf{r} \frac{\partial \mathbf{T}}{\partial \mathbf{r}} \right) + \frac{\partial^2 \mathbf{T}}{\partial \mathbf{z}^2} \right] + \frac{2\mathbf{v}}{c_p} \left[\left(\frac{\partial \mathbf{u}}{\partial \mathbf{r}} \right)^2 + \left(\frac{\partial \mathbf{w}}{\partial \mathbf{z}} \right)^2 \right] \\ + \frac{\mathbf{v}}{c_p} \left[\left(\frac{\partial \mathbf{v}}{\partial \mathbf{z}} \right)^2 + \left(\frac{\partial \mathbf{w}}{\partial \mathbf{r}} + \frac{\partial \mathbf{u}}{\partial \mathbf{z}} \right)^2 + \left\{ \mathbf{r} \frac{\partial}{\partial \mathbf{r}} \left(\frac{\mathbf{v}}{\mathbf{r}} \right)^2 \right]$$
(3)

The velocity boundary conditions are:

$$u(r,0) = 0 \qquad u(r,\infty) = 0$$
$$v(r,0) = r\omega \qquad v(r,\infty) = 0 \qquad (4)$$
$$w(r,0) = w_{rv}$$

In the above, w_w is the fluid velocity at the wall. Positive values of w_w correspond to blowing, and negative values correspond to suction.

For time $t \leq 0$, there will be a steady temperature field in the fluid due to frictional dissipation. Its precise nature would depend on conditions prevailing in the disk. As an example, one may consider a disk of infinite thermal conductivity with the consequence that its temperature rise due to the frictional dissipation is uniform. On the other hand, the disk may have large resistance to heat flow in the radial direction. This situation contrasts with that for flow over a flat plate. With these observations, we consider two cases of simple initial and boundary conditions for the temperature field as follows:

Case (i) For disks initially at adiabatic wall temperature.

$$T(r, z, 0) = T_{i,1}(r, z)$$
$$- \frac{\partial T(r, 0, t)}{\partial z} = \frac{q_{W}}{k} 1(t)$$
$$T(r, \infty, 0) = T_{\infty}$$
$$\frac{\partial T(r, 0, 0)}{\partial z} = 0$$

In the above $T_{i,1}(r,z)$ represents the initial temperature distribution in the fluid consistent with an adiabatic wall temperature T_{aw} . Here the problem is to examine the transient response of the wall temperature due to a step change in wall heat flux.

4

(5.a)

Case (ii) For disks initially of uniform wall temperature

$$T(r,z,0) = T_{i,2}(r,z)$$

-
$$\frac{\partial T(r,0,t)}{\partial z} = \frac{q_{iw}(r)}{k} + \frac{1}{k} \left[q_w(r) - q_{iw}(r) \right] 1(t)$$

$$T(r,w,t) = T_{-}$$

 $T(r,0,0) = T_{iw} = constant$

Where q_{iw}(r) is the wall flux due to frictional dissipation.

The initial temperature field $T_{i,2}$ is clearly different from that in Case (i). It is to be noted that for a disk of uniform temperature, the wall flux q_{iw} is not uniform. Here we are examining the transient response of the wall temperature due to a <u>uniform</u>, step change of the local wall flux. While $q_w(r)$ and $q_{iw}(r)$ both vary with r, their difference is a constant being independent of the radius.

3.0 Method of Solution

The solution of (1), (2), and (3) with the initial and the boundary conditions (4) and (5) for moderate suction is given in [9]. In this report we consider large values of suction. Physical considerations suggest expressing the velocity components u and v as:

 $u = r \otimes F(\eta) , \qquad v = r \otimes G(\eta)$ (6.a)
wherein $\eta = z \left(\frac{\omega}{v}\right)^{1/2}$.

It follows, then, from the continuity requirement that

$$\mu = (\psi \ v) H(\eta)$$
(6.b)

Upon introducing the foregoing expressions for the velocity components into the momentum equations and observing that the pressure must necessarily approach a constant value at infinity, one is led to the conclusion that P is independent of r and may be expressed as

$$\mathbf{p} = \rho \, \mathbf{w} \, \mathbf{v} \, \mathbf{P}(\eta)$$

(6.c)

(5.b)

Substituting (6.a, b, and c) into (1) and (2.a, b, and c) yields, after some rearrangement,

$$G'' = HG' - H'G$$
 (7.a)*

 $H'' = HH'' - \frac{1}{2} (H')^2 + 2G^2$ (7.b)*

(7.c)*

(8)

and $F = -\frac{H}{2}$

with

$$G(0) = 1, \qquad H(0) = -a_0 = \frac{w_W}{(W \ v)^{1/2}}, \qquad H'(0) = 0$$

$$G(\infty) = 0, \qquad H'(\infty) = 0$$

In the foregoing, the prime denotes differentiation with respect to η ; a_0 is the mass transfer parameter. Numerical solutions of the above set of equations for an impervious wall, i.e. $a_0 = 0$, were given by Cochran [14]. Stuart presented an analytical solution for large suction. His result will be used later for the integration of the unsteady energy equation in the present analysis.

Sparrow and Gregg [11] reported computer results for a ranging from -5 to 4.

To transform the energy equation, we further let

. ...

$$R = r \left(\frac{\omega}{v}\right)^{1/2} , \qquad \tau = \frac{\omega t}{P_r}$$
(9)

and obtain

$$\frac{\partial \mathbf{T}}{\partial \tau} + \mathbf{P}_{\mathbf{r}} \cdot \mathbf{R} \cdot \mathbf{F} \frac{\partial \mathbf{T}}{\partial \mathbf{R}} + \mathbf{P}_{\mathbf{r}} \cdot \mathbf{H} \cdot \frac{\partial \mathbf{T}}{\partial \eta} = \frac{1}{\mathbf{R}} \frac{\partial}{\partial \mathbf{R}} \left(\mathbf{R} \cdot \frac{\partial \mathbf{T}}{\partial \mathbf{R}} \right) + \frac{\partial^2 \mathbf{T}}{\partial \eta^2}$$

$$+ \frac{P_{r} w v}{C_{p}} R^{2} \left[(G')^{2} + (F')^{2} \right] + \frac{2 P_{r} w v}{C_{p}} \left[2F^{2} + (H')^{2} \right]$$
(10)

A fourth equation which governs the pressure distribution in the fluid is not of our concern and is thus omitted.

We shall first consider the case for the adiabatic wall temperature with initial and boundary conditions prescribed by (5.a). For convenience, we introduce a dimensionless temperature defined by

$$\frac{\mathbf{T} - \mathbf{T}_{\infty}}{\frac{qw}{k} \left(\frac{v}{w}\right)^{1/2}} = \frac{w v}{\mathbf{c}_{p} \frac{qw}{k} \left(\frac{v}{w}\right)^{1/2}} \quad (\mathbf{R}^{2}\mathbf{S} + \mathbf{Q}) + \theta(\eta, \tau)$$
(11)

The functions $S(\eta)$, $Q(\eta)$, and $\theta(\eta, \tau)$ respectively, satisfy

S'' - P_r . H . S' + P_r . H' . S = - P_r
$$[(G')^2 + (F')^2]$$
 (12.a)

S'(0) = 0, $S(\infty) = 0$ $Q'' - P_r \cdot H \cdot Q' = -4 (S + 3 P_r \cdot F^2)$ (13.a)

(12.b)

with
$$Q'(0) = 0$$
, $Q(\infty) = 0$ (13.b)

and

$$\frac{\partial \theta}{\partial \tau} + \mathbf{P}_{\mathbf{r}} \cdot \mathbf{H} \frac{\partial \theta}{\partial \eta} = \frac{\partial^2 \theta}{\partial \eta^2}$$
 (14.a)

with
$$\theta(\eta,0) = 0$$
 $\frac{\partial \theta(0,\tau)}{\partial \eta} = -1(\tau), \quad \theta(\infty,\tau) = 0$ (14.b)

For disks initially at the uniform wall temperature, T subjected to a uniform, step change in wall flux, we let

$$\frac{T - T_{\infty}}{\left(\frac{\nu}{\omega}\right)^{1/2} \frac{\Delta q_{w}}{k}} = \frac{\omega \nu}{c_{p} \frac{\Delta q_{w}}{k} \left(\frac{\nu}{\omega}\right)^{1/2}} \left[R^{2} S(\eta) + Q(\eta) \right] + \theta(\eta, \tau)$$
(15)

Here $\triangle q_w = q_w(R) - q_{iw}(R)$ which is a constant. The functions S and Q, respectively, satisfy (12.a) and (13.a) but with boundary conditions altered to

> S(0) = 0 , $S(\infty) = 0$ (16.a)

and

$$Q(0) = \frac{cp}{\omega} (T_{i\omega} - T_{\omega}) , \qquad Q(\omega) = 0 \qquad (16.b)$$

It is easy to show that θ of Eqn. 15 satisfies Eqn. 14.a with initial and boundary conditions of Eqn. 14.b.

Finally, we note that for negligible dissipation the fluid will initially have a uniform temperature T_{∞} , which is also the wall temperature prior to any thermal disturbance. If we now re-define $\theta(\eta, \tau)$ according to

$$= \frac{\mathbf{T} - \mathbf{T}_{\infty}}{\frac{\mathbf{q}_{w}(\mathbf{v})}{\mathbf{k}}^{1/2}}$$

it is easy to demonstrate that θ again satisfies (14.a) with initial and boundary conditions (14.b).

To integrate (14.a) with (14.b) for large values of the suction parameter (say $a_0 = 1.5$ or greater), we employ the Stuart's solution [15] for the velocity field in inverse powers of a_0 . In particular, the H function is given as,

$$H(\xi) = -a_{0} + \frac{1}{a_{0}^{3}} \left(-\frac{1}{2} + e^{-\xi} - \frac{1}{2} e^{-2\xi} \right)$$

+ $\frac{1}{a_{0}^{7}} \left[\frac{201}{288} + \left(\frac{-1}{2} \xi - \frac{95}{72} \right) e^{-\xi} + \left(\frac{1}{2} \xi + \frac{13}{24} \right) e^{-2\xi} + \frac{1}{12} e^{-3\xi} - \frac{1}{288} e^{-4\xi} \right] + 0 (a_{0}^{-11})$ (17)

wherein $\xi = a \eta$

Application of Laplace transform to (14.a) with initial and boundary conditions (14.b) yields

$$\theta'' - P_r \cdot H \cdot \theta' = p\theta$$
 (18)

with $\theta'(0) = -\frac{1}{p}$ and $\theta(\infty) = 0$. We now introduce a new variable

Y(n,p) defined by

$$\bar{\mathbf{Y}} = \mathbf{p}\bar{\boldsymbol{\theta}} \exp\left(-\frac{\mathbf{P}_{\mathbf{r}}}{2}\int_{0}^{\eta}\mathbf{H}(\eta)\,\mathrm{d}\eta\right)$$
(19)

and obtain from (18)

$$\frac{1}{Y''} + \left(-p + \frac{P_r}{2} H' - \frac{P_r^2}{4} H^2\right) \bar{Y} = 0$$

with

$$-\frac{a_{o}P}{2}Y(0) + a_{o}Y'(0) = -1, \quad Y(\infty) = 0$$

In (18) and (20) the prime denotes differentiation with respect η . Transforming the independent variable from η to ξ and making use of (17), for large a_0 , we arrive at,

$$\frac{d^2 \Upsilon(5)}{d5^2} - \left(c_0^2 + c_1 e^{-5} + c_2 e^{-25} + c_3 e^{-35} + \dots\right) \tilde{\Upsilon}(5) = 0$$
(21)

wherein

$$c_{0}^{2} = \frac{p}{a_{0}^{2}} + \frac{\frac{p}{r}}{4} \left[1 + \frac{1}{a_{0}^{4}} + 0\left(\frac{1}{a_{0}^{8}}\right) \right], c_{1} = \frac{\frac{p}{r}(1 - \frac{p}{r})}{2 a_{0}^{4}} \left[1 + 0\left(\frac{1}{a_{0}^{4}}\right) \right]$$

$$c_{2} = \frac{\frac{p}{r}\left(\frac{p}{r} - 2\right)}{4a_{0}^{4}} \left[1 + 0\left(\frac{1}{a_{0}^{4}}\right) \right], c_{3} = \frac{\frac{p}{r}\left(1 - \frac{7}{3}\frac{p}{r}\right)}{8a_{0}^{8}} \left[1 + 0\left(\frac{1}{a_{0}^{4}}\right) \right] etc.$$

The solution of (21) satisfying the boundary condition

 $-\frac{a_{o}P}{2} Y(0) + a_{o} Y'(0) = -1 \text{ and } Y(\infty) = 0 \text{ may be appropriately represented}$ by a series of the form

$$\overline{Y} = Ke^{-C_0^5} (1 + A_1 e^{-5} + A_2 e^{-25} + A_3 e^{-35} + \dots)$$
(22)

(20)

where

$$K = \frac{1}{a_0 \left[\left(\frac{P_r}{2} + c_0 \right) \left(1 + A_1 + A_2 + A_3 + \dots \right) + \left(A_1 + 2A_2 + 3A_3 + \dots \right) \right]}$$

$$A_1 = \frac{c_1}{1 + 2c_0} , \qquad A_2 = \frac{c_1^2 + c_2(1 + 2 c_0)}{4 (1 + 2 c_0) (1 + c_0)}$$

$$A_3 = \frac{c_1^3 + c_1 c_2(5 + 6 c_0) + 4c_3(1 + 2 c_0) (1 + c_0)}{12 (1 + 2 c_0) (1 + c_0) (3 + 3 c_0)} , \text{ etc.}$$

The latter are obtained by substituting (22) into (21) and equating coefficients of terms like $e^{-(C_0 + 1)\xi}$, $e^{-(C_0 + 2)\xi}$..., to zero. Since $C_1 = 0$ $(\frac{1}{a_0^4})$, $C_2 = 0$ $(\frac{1}{a_0^4})$, but $C_3 = 0$ $(\frac{1}{a_0^8})$, it is a valid

approximation to ignore terms involving C_1^2 , C_1C_2 , C_3 , and C_1^3 . Accordingly, after substituting the values of the A_n , K may also be written as

$$\kappa = \frac{\frac{1}{2a_0} \left(1 + 2 c_0\right) \left(1 + c_0\right)}{c_0^3 + \left(\frac{3}{2} + \frac{c_1}{2} + \frac{c_2}{4} + \frac{p_r}{2}\right) c_0^2 + \left(\frac{1}{2} + c_1 + \frac{5}{8}c_2 + \frac{3p_r}{4} + \frac{c_1^2 r}{4} + \frac{c_2^2 r}{8}\right) c_0}$$

 $+\left(\frac{c_1}{2}+\frac{c_2}{4}\right) + \frac{P_r}{4}\left(1+c_1+\frac{c_2}{4}\right)$ (Note denominator only continued to this line)

(23)

and thus (22) may be approximated by

$$\overline{Y} = e^{-C_0 5} \left[\frac{C_0^2 + C_0 \left(\frac{3}{2} + \frac{C_1}{2} e^{-5} + \frac{C_2}{4} e^{-25}\right) + \frac{1}{2} + \frac{C_1}{2} e^{-5} + \frac{C_2}{8} e^{-25}}{a_0 (C_0 - \Lambda_1) (C_0 - \Lambda_2) (C_0 - \Lambda_3)} \right]$$
(24)

where Λ_1 ; Λ_2 and Λ_3 are the three roots of the cubic equation in the denominator of (23). These roots are characterized by the determinant of this cubic equation. For convenience, let

 $p_{1} = \frac{3}{2} + \frac{C_{1}}{2} + \frac{C_{2}}{4} + \frac{P_{r}}{2} \qquad q = \frac{1}{2} + C_{1} + \frac{5C_{2}}{8} + \frac{3P_{r}}{4} + \frac{C_{1}P_{r}}{4} + \frac{C_{2}P_{r}}{4}$ $r = \frac{C_{1}}{2} + \frac{C_{2}}{4} + \frac{P_{r}}{4} \left(1 + C_{1} + \frac{C_{2}}{4}\right)$ $a = \frac{1}{3} \left(3q - p_{1}^{2}\right) \qquad b = \frac{1}{27} \left(2 p_{1}^{3} - 9qp_{1} + 27 r\right)$

Then if $\frac{b^2}{4} + \frac{a^3}{27} > 0$, one real and two complex conjugate roots will be obtained as follows;

let
$$A = \left[\frac{b}{2} + \left(\frac{b^2}{4} + \frac{a^3}{27}\right)^{1/2}\right]^{1/3}$$
 $B = \left[-\frac{b}{2} - \left(\frac{b^2}{4} + \frac{a^3}{27}\right)^{1/2}\right]^{1/3}$

we may write,

$$\Lambda_{1} = \beta_{1} = A + B - \frac{p_{1}}{3}$$
$$\Lambda_{2} = \beta_{2} - i\beta_{3} = -\frac{A + B}{2} + \frac{A - B}{2}\sqrt{-3} - \frac{p_{1}}{3}$$
$$\Lambda_{3} = \beta_{2} + i\beta_{3} = -\frac{A + B}{2} - \frac{A - B}{2}\sqrt{-3} - \frac{p_{1}}{3}$$

In the above, β_1 is the real root and β_2 is the real part and the β_3 is the imaginary part of the two conjugate roots.

let

If $\frac{b^2}{4} + \frac{a^3}{27} < 0$, three unequal real roots will be obtained as follows:

$$\cos \varphi = -\frac{b}{2} / \sqrt{-\frac{a}{27}}$$

$$A_{1} = 2\sqrt{\frac{-a}{3}} \cos \frac{\varphi}{3} - \frac{p_{1}}{3}$$

$$A_{2} = 2\sqrt{\frac{-a}{3}} \cos \left(\frac{\varphi}{3} + \frac{2\pi}{3}\right) - \frac{p_{1}}{3}$$

$$A_{3} = 2\sqrt{\frac{-a}{3}} \cos \left(\frac{\varphi}{3} + \frac{4\pi}{3}\right) - \frac{p_{1}}{3}$$

If $\frac{b^2}{4} + \frac{a^3}{27} = 0$, three real roots will be obtained but at least two are equal.

In view of the foregoing analysis, we substitute (24) into (19); revert back to the η variable, consistently ignore terms of small order of magnitude, take the inverse transform and obtain for

$$\frac{b^2}{4} + \frac{a^3}{27} > 0$$
 and $\alpha \neq \alpha_0^2 \beta_1^2$

the dimensionless temperature function as:

$$\theta(\eta,\tau) = \exp\left[-\frac{P_{r}}{2}\left(a_{0} + \frac{P_{r}}{2a_{0}^{3}}\right)\eta - \frac{P_{r}}{2a_{0}^{4}}\left(e^{-a_{0}\eta} - \frac{1}{4}e^{-2a_{0}\eta} - \frac{3}{4}\right)\right]$$

$$\times \left\{ \frac{\Psi_1}{2} \left[\frac{1}{-a_0 \beta_1 + \alpha^{1/2}} e^{-\alpha^{1/2} \eta} \operatorname{erfc} \left[\frac{\eta}{2\alpha^{1/2}} - (\alpha \tau)^{1/2} \right] \right\}$$

$$+ \frac{e^{\alpha^{1/2}\eta}}{-a_0^{\beta_1} - \alpha^{1/2}} \operatorname{erfc} \left[\frac{\eta}{2\tau^{1/2}} + (\alpha \tau)^{1/2} \right]$$

(continued next page)

$$\begin{aligned} &+ \frac{\pi}{a_{0}^{2}\beta_{1}^{2} - \alpha}{2} \exp\left[-a_{0}\beta_{1}^{\eta} + (a_{0}^{2}\beta_{1}^{2} - \alpha)^{\gamma}\right] \operatorname{erfc}\left(\frac{\eta}{2\tau^{1/2}} - a_{0}\beta_{1}\tau^{1/2}\right) \\ &+ 2\operatorname{Re}\left(\frac{\pi}{2^{2}}\left[\frac{e^{-\alpha^{1/2}\eta}}{-a_{0}\left(+\beta_{2} - i\beta_{3}\right) + \alpha^{1/2}} - \operatorname{erfc}\left[\frac{\eta}{2\tau^{1/2}} - (\alpha \tau)^{1/2}\right]\right] \\ &+ \frac{e^{-\alpha\tau}}{a_{0}\left(-\beta_{2} + i\beta_{3}\right) - \alpha^{1/2}} - \operatorname{erfc}\left[\frac{\eta}{2\tau^{1/2}} + (\alpha \tau)^{1/2}\right] \\ &+ e^{-\alpha\tau} \cdot \frac{\pi}{2a_{0}\left(\beta_{2} - i\beta_{3}\right)}{a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2} - \alpha} \exp\left[-a_{0}\left(\beta_{2} - i\beta_{3}\right)\eta + a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2}\eta\right] \\ &+ e^{-\alpha\tau} \cdot \frac{\pi}{2a_{0}\left(\beta_{2} - i\beta_{3}\right)^{2} - \alpha} \exp\left[-a_{0}\left(\beta_{2} - i\beta_{3}\right)\eta + a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2}\eta\right] \\ &+ e^{-\alpha\tau} \cdot \frac{\pi}{2a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2} - \alpha} \exp\left[-a_{0}\left(\beta_{2} - i\beta_{3}\right)\eta + a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2}\eta\right] \\ &+ e^{-\alpha\tau} \cdot \frac{\pi}{2a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2} - \alpha} \exp\left[-a_{0}\left(\beta_{2} - i\beta_{3}\right)\eta + a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2}\eta\right] \\ &+ e^{-\alpha\tau} \cdot \frac{\pi}{2a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2} - \alpha} \exp\left[-a_{0}\left(\beta_{2} - i\beta_{3}\right)\eta + a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2}\eta\right] \\ &+ e^{-\alpha\tau} \cdot \frac{\pi}{2a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2} - \alpha} \exp\left[-a_{0}\left(\beta_{2} - i\beta_{3}\right)\eta + a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2}\eta\right] \\ &+ e^{-\alpha\tau} \cdot \frac{\pi}{2a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2} - \alpha} \exp\left[-a_{0}\left(\beta_{2} - i\beta_{3}\right)\eta + a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2}\eta\right] \\ &+ e^{-\alpha\tau} \cdot \frac{\pi}{2a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2} - \alpha} \exp\left[-a_{0}\left(\beta_{2} - i\beta_{3}\right)\eta + a_{0}^{2}\left(\beta_{2} - i\beta_{3}\right)^{2}\eta\right] \\ &+ e^{-\alpha\tau} \left[-\frac{\pi}{2\tau^{1/2}} - a_{0}\left(\beta_{2} - i\beta_{3}\right)\eta + \frac{\pi}{2a_{0}^{2}\left(\alpha\tau} - \alpha^{1/2}\right) + \frac{\pi}{2a_{0}^{2}\left(\alpha\tau^{2} - i\beta_{3}\right)^{2}\eta} \right] \\ &+ e^{-\alpha\tau} \left[-\frac{\pi}{2\tau^{1/2}} + \frac{\pi}{2a_{0}^{3}}\right] \left[-\frac{\pi}{2a_{0}^{3}} + \frac{\pi}{2a_{0}^{3}}\right] \\ &+ e^{-\alpha\tau} \left[-\frac{\pi}{2\tau^{1/2}} + \frac{\pi}{2a_{0}^{3}}\right] \left[-\frac{\pi}{2a_{0}^{3}} + \frac{\pi}{2a_{0}^{3}}\right] \\ &+ e^{-\alpha\tau} \left[-\frac{\pi}{2\tau^{1/2}} + \frac{\pi}{2\tau^{1/2}}\right] \\ &+ e^{-\alpha\tau^{1/2}\eta} \exp\left[-\frac{\pi}{2\tau^{1/2}} + \frac{\pi}{2\tau^{1/2}}\right] \\ &+ e^{-\alpha\tau^{$$

+ 2 Re (---- Repeat 25.a) (25.b)

The quantity in the bracket after 2Re in the equation (25.b) is same as that in equation (25.a). In the above the symbols α , Ψ_1 and Ψ_2 are defined as:

$$\alpha = \frac{\Pr_{r}^{2} a_{0}^{2}}{4} \left(1 + \frac{1}{a_{0}^{4}}\right)$$

$$\Psi_{1}(\eta) = \frac{\beta_{1}^{2} + \frac{1}{2} (3\beta_{1} + 1) + \frac{\Pr_{r}(1 - \Pr_{r})}{4a_{0}^{4}} (\beta_{1} + 1)e^{-a_{0}\eta} + \frac{\Pr_{r}(\Pr_{r} - 2)}{16a_{0}^{4}} (\beta_{1} + \frac{1}{2})e^{-2a_{0}\eta}}{(\beta_{2} - \beta_{1})^{2} + \beta_{3}^{2}}$$

$$\Psi_{2}(\eta) = \frac{\left[\left(\beta_{2} - \beta_{1}\right)^{1} - \beta_{3}\right]\left[(\beta_{2} - i\beta_{3})^{2} + \frac{3}{2}(\beta_{2} - i\beta_{3}) + \frac{1}{2} + \frac{\Pr_{r}(1 - \Pr_{r})}{4a_{0}^{4}} (\beta_{2} - i\beta_{3} + 1)e^{a_{0}\eta}}{2\beta_{3}\left[(\beta_{1} - \beta_{2})^{2} + \beta_{3}^{2}\right]}$$

$$P_{2}(\eta) = \frac{\Pr_{r}(\Pr_{r} - 2)}{2\beta_{3}\left[(\beta_{1} - \beta_{2})^{2} + \beta_{3}^{2}\right]}$$

$$+\frac{P_{r}(P_{r}-2)}{16a_{0}}(\beta_{2}-i\beta_{3}+\frac{1}{2})\bar{e}^{2a_{0}}\eta]$$

(Numerator only continued to this line)

for $\frac{b^2}{4} + \frac{a^3}{27} < 0$, and for $\alpha \neq a_0^2 \Lambda_1^2$; $\alpha \neq a_0^2 \Lambda_2^2$; $\alpha \neq a_0^2 \Lambda_3^2$ the dimensionless temperature θ takes the following form:

$$\theta(\eta,\tau) = \exp\left[-\frac{P_r}{2}\left(a_0 + \frac{P_r}{2a_0^3}\right)\eta - \frac{P_r}{2a_0^4}\left(e^{-a_0\eta} - \frac{1}{4}e^{-2a_0\eta} - \frac{3}{4}\right)\right]$$

$$\left\{\frac{1}{2}\left[\frac{\Phi_{1}(\eta)}{\alpha^{1/2}}+\frac{\Phi_{2}(\eta)}{\alpha^{1/2}}+\frac{\Phi_{3}(\eta)}{\alpha^{1/2}}+\frac{\Phi_{3}(\eta)}{\alpha^{1/2}}\right]e^{-\alpha^{1/2}\eta} e^{-\alpha^{1/2}\eta} e^{-\alpha^{1/2}\eta} e^{-\alpha^{1/2}\eta}\right\}$$

$$\frac{1}{2} \left[\frac{\Phi_1(\eta)}{\alpha^{1/2} + a_0 \Lambda_1} + \frac{\Phi_2(\eta)}{\alpha^{1/2} + a_0 \Lambda_2} + \frac{\Phi_3(\eta)}{\alpha^{1/2} + a_0 \Lambda_3} \right] e^{\alpha^{1/2} \eta} \operatorname{erfq} \left[\frac{\eta}{2\tau^{1/2}} + (\alpha\tau)^{1/2} \right]$$

(continued next page)

$$+ a_{0}e^{-\alpha \tau} \left[\frac{1}{a_{0}^{2}\Lambda_{1}^{2} - \alpha} \exp\left(-a_{0}\Lambda_{1}^{\eta} + a_{0}^{2}\Lambda_{1}^{2\gamma}\right) \operatorname{erfc}\left(\frac{\eta}{2\tau^{1/2}} - a_{0}\Lambda_{1}^{\tau^{1/2}}\right) \right] + \frac{\Lambda_{2}^{\frac{5}{2}}(\eta)}{a_{0}^{2}\Lambda_{2}^{2} - \alpha} \exp\left(-a_{0}\Lambda_{2}^{\eta} + a_{0}^{2}\Lambda_{2}^{2\gamma}\right) \operatorname{erfc}\left(\frac{\eta}{2\tau^{1/2}} - a_{0}\Lambda_{2}^{\tau^{1/2}}\right) + \frac{\Lambda_{3}^{\frac{5}{2}}(\eta)}{a_{0}^{2}\Lambda_{2}^{2} - \alpha} \exp\left(-a_{0}\Lambda_{3}^{\eta} + a_{0}^{2}\Lambda_{3}^{2\gamma}\right) \operatorname{erfc}\left(\frac{\eta}{2\tau^{1/2}} - a_{0}\Lambda_{3}^{\tau^{1/2}}\right) \right] \right\}$$
(25.c)

also for $\frac{b^2}{4} + \frac{a^3}{27} < 0$, and $\alpha = a_0^2 \Lambda_1^2$ which automatically implies $\alpha \neq a_0^2 \Lambda_2^2$ and $\alpha \neq a_0^2 \Lambda_3^2$ the temperature function is expressed as:

$$\theta(\eta,\tau) = \exp\left[-\frac{P_{r}}{2}\left(a_{0} + \frac{P_{r}}{2a_{0}^{3}}\right)\eta - \frac{P_{r}}{2a_{0}^{4}}\left(e^{-a_{0}\eta} - \frac{1}{4}e^{-2a_{0}\eta} - \frac{3}{4}\right)\right]$$

 $\alpha = \Gamma \Lambda_1 \Phi_1(\eta)$

$$\left\{ \left[\frac{\frac{\Phi}{2}}{4\alpha^{1/2}} + \frac{\frac{\Phi}{2}}{2(\alpha^{1/2} - a_0 \Lambda_2)} + \frac{\frac{\Phi}{3}}{2(\alpha^{1/2} - a_0 \Lambda_3)} \right] e^{-\alpha^{1/2} \eta} \operatorname{erfd}\left[\frac{\eta}{2\tau^{1/2}} - (\alpha\tau)^{1/2} \right] \right\}$$

$$-\left[\frac{\frac{\Phi}{2}}{2\alpha^{1/2}}\left(\frac{1}{2}+\alpha^{1/2}\eta+2\alpha\tau\right)+\frac{\frac{\Phi}{2}}{2(\alpha^{1/2}+a_{0}\Lambda_{2})}+\frac{\frac{\Phi}{3}}{2(\alpha^{1/2}+a_{0}\Lambda_{3})}\right]e^{\alpha^{1/2}\eta}$$

$$\operatorname{erfc}_{2\tau^{1/2}}^{\left[\frac{1}{2}\right]} + \left[\left(\alpha\tau\right)^{1/2}\right] + \left[\left(\frac{\tau}{4\tau}\right)^{1/2}\right] \exp\left(-\frac{\eta^2}{4\tau} - \alpha\tau\right) \quad (eqno(1))$$

equation continued on next page)

When $\alpha = a_0^2 \Lambda_2^2$ (But $\alpha \neq a_0^2 \Lambda_1^2$), interchange Φ_1 and Φ_2 and replace Λ_2 by Λ_1 . Similarly rule may be applied when $\alpha = a_0^2 \Lambda_3^2$.

$$+ \frac{a_{0}A_{2}^{\Phi}}{a_{0}^{2}A_{2}^{2} - \alpha} \exp\left[\left(a_{0}^{2}A_{2}^{2} - \alpha\right)\tau - a_{0}A_{2}\eta\right] \operatorname{erfc}\left(\frac{\eta}{2\tau^{1/2}} - a_{0}A_{2}\tau^{1/2}\right) + \frac{a_{0}A_{3}^{\Phi}}{a_{0}^{2}A_{3} - \alpha} \exp\left[\left(a_{0}^{2}A_{3}^{2} - \alpha\right)\tau - a_{0}A_{3}\eta\right] \operatorname{erfc}\left(\frac{\eta}{2\tau^{1/2}} - a_{0}A_{3}\tau^{1/2}\right)\right]$$

where
$$\Lambda_{1}^{2} + \frac{3\Lambda}{2} + \frac{1}{2} + \frac{P_{r}(1 - P_{r})}{4a_{0}^{4}} (1 + \Lambda_{1})e^{-a_{0}\eta} + \frac{P_{r}(P_{r} - 2)}{16a_{0}^{4}} (\Lambda_{1} + \frac{1}{2})e^{-2a_{0}\eta}$$

 $\Phi_{1}(\eta) = \frac{(\Lambda_{1} - \Lambda_{2})(\Lambda_{1} - \Lambda_{3})}{(\Lambda_{1} - \Lambda_{3})}$
(25.d)

$$\Phi_{2} = \frac{\Lambda_{2}^{2} + \frac{3}{2}\Lambda_{2} + \frac{1}{2} + \frac{P_{r}(1 - P_{r})}{4a_{0}^{4}}(1 + \Lambda_{2})e^{-a_{0}\eta} + \frac{P_{r}(P_{r} - 2)}{16a_{0}^{4}}(\Lambda_{2} + \frac{1}{2})e^{-2a_{0}\eta}}{(\Lambda_{2} - \Lambda_{1})(\Lambda_{2} - \Lambda_{3})}$$

$$\Phi_{3}^{2} + \frac{3}{2} \Lambda_{3} + \frac{1}{2} + \frac{P_{r}(1 - P_{r})}{4a_{0}^{4}} (1 + \Lambda_{3})e^{-a_{0}^{T}} + \frac{P_{r}(P_{r} - 2)}{16a_{0}^{4}} (\Lambda_{3} + \frac{1}{2})e^{-2a_{0}^{T}}$$

$$(\Lambda_{3} - \Lambda_{1}) (\Lambda_{3} - \Lambda_{2})$$

for
$$\frac{b^2}{4} + \frac{a^3}{27} = 0$$
, and for $\alpha \neq a_0^2 \Lambda_1^2$, and $\alpha \neq a_0^2 \Lambda_2^2$

the function θ is given as:

$$\theta(\eta,\tau) = \exp\left[-\frac{P_r}{2}\left(a_0 + \frac{P_r}{2a_0^3}\right)\eta - \frac{P_r}{2a_0^4}\left(e^{-a_0\eta} - \frac{1}{4}e^{-a_0\eta} - \frac{3}{4}\right)\right]$$

$$\left[\frac{\chi_{1}a_{0}}{2\alpha^{1/2}-a_{0}\Lambda_{2}}^{2}+\frac{\chi_{2}}{2\alpha^{1/2}-a_{0}\Lambda_{2}}^{2}+\frac{\chi_{3}}{2(\alpha^{1/2}-a_{0}\Lambda_{2})}^{2}+\frac{\chi_{3}}{2(\alpha^{1/2}-a_{0}\Lambda_{2})}^{2}\right]e^{-\alpha^{1/2}\eta}\operatorname{erfc}\left(\frac{\eta}{2\tau^{1/2}}-(\alpha\tau)^{1/2}\right)$$

$$+ \left[\frac{\chi_{1}a_{0}}{\chi(x^{1/2} - a_{0}A_{2})^{2}} - \frac{\chi_{2}}{\chi(x^{1/2} + a_{0}A_{2})} - \frac{\chi_{3}}{\chi(x^{1/2} + a_{0}A_{1})} \right] e^{\alpha^{1/2}\eta} \operatorname{erfd}\left[\frac{\eta}{2\tau^{1/2}} + (\alpha\tau)^{1/2} \right]$$

$$+ \left[\chi_{1}a_{0} \left(\frac{1 - a_{0}A_{2}\eta + 2a_{0}^{2}A_{2}^{2}\tau}{a_{0}^{2}A_{2}^{2} - \alpha} - \frac{a_{0}A_{2}}{(a_{0}^{2}A_{2}^{2} - \alpha)^{2}} \right) + \frac{\chi_{2}a_{0}A_{2}}{(a_{0}^{2}A_{2}^{2} - \alpha)} \right] \cdot$$

$$\exp\left(-a_{0}A_{2}\eta + (a_{0}^{2}A_{2}^{2} - \alpha) \tau \right) \operatorname{erfd}\left(\frac{\eta}{2\tau^{1/2}} - a_{0}A_{2}\tau^{1/2} \right) +$$

$$+ \frac{\chi_{4}a_{0}}{\pi^{1/2}} \exp\left(- \frac{\eta^{2}}{4\tau} - \alpha\tau \right) \left[\frac{2a_{0}A_{2}}{a_{0}^{2}A_{2}^{2} - \alpha} - \frac{1}{(a_{0}^{2}A_{2}^{2} - \alpha)^{2} - \tau^{1/2}} \right]$$

$$+ \frac{\chi_{4}a_{0}}{\pi^{1/2}} \exp\left(- \frac{\eta^{2}}{4\tau} - \alpha\tau \right) \left[\frac{2a_{0}A_{2}}{a_{0}^{2}A_{2}^{2} - \alpha} - \frac{1}{(a_{0}^{2}A_{2}^{2} - \alpha)^{2} - \tau^{1/2}} \right]$$

$$+ \frac{\chi_{4}a_{0}}{(a_{0}^{2}A_{1}^{2} - \alpha)} \exp\left(- a_{0}A_{1}\eta + (a_{0}^{2}A_{1}^{2} - \alpha) \tau \right) \operatorname{erfd}\left(\frac{\eta}{2\tau^{1/2}} - a_{0}A_{1}\tau^{1/2} \right)$$

$$+ \frac{\chi_{4}a_{0}}{(a_{0}^{2}A_{1}^{2} - \alpha)} \exp\left(- a_{0}A_{1}\eta + (a_{0}^{2}A_{1}^{2} - \alpha) \tau \right) \operatorname{erfd}\left(\frac{\eta}{2\tau^{1/2}} - a_{0}A_{1}\tau^{1/2} \right)$$

$$+ \frac{\chi_{4}a_{0}}{(a_{0}^{2}A_{1}^{2} - \alpha)} \exp\left(- a_{0}A_{1}\eta + (a_{0}^{2}A_{1}^{2} - \alpha) \tau \right) \operatorname{erfd}\left(\frac{\eta}{2\tau^{1/2}} - a_{0}A_{1}\tau^{1/2} \right)$$

$$+ \frac{\chi_{4}a_{0}A_{1}}{(a_{0}^{2}A_{1}^{2} - \alpha)} \exp\left(- a_{0}A_{1}\eta + (a_{0}^{2}A_{1}^{2} - \alpha) \tau \right) \operatorname{erfd}\left(\frac{\eta}{2\tau^{1/2}} - a_{0}A_{1}\tau^{1/2} \right)$$

$$+ \frac{\chi_{4}a_{0}A_{1}}{(a_{0}^{2}A_{1}^{2} - \alpha)} \exp\left(- \frac{\eta}{2a_{0}^{2}} + \frac{\chi_{2}}{2a_{0}^{2}} \right) \eta - \frac{\eta}{2a_{0}} \left\{ \left(e^{-a_{0}\eta} - \frac{1}{4} e^{-2a_{0}\eta} - \frac{\eta}{2} \right) \right]$$

$$+ \left[\left(\frac{\chi_{4}a_{0}}{\chi(\alpha^{1/2} - a_{0}A_{2}} \right)^{2} + \frac{\chi_{2}}{\chi(\alpha^{1/2} - a_{0}A_{2}} \right) + \frac{\chi_{3}}{4\alpha^{1/2}} \right] e^{-\alpha^{1/2}\eta} \operatorname{erfd}\left[\frac{\eta}{2\tau^{1/2}} - (\alpha\tau)^{1/2} \right]$$

$$\begin{aligned} &+ \left[\frac{\chi_{1}a_{0}}{2(\alpha^{1/2} - a_{0}\Delta_{2})^{2}} - \frac{\chi_{2}}{2(\alpha^{1/2} + a_{0}\Delta_{2})} - \frac{\chi_{3}}{2\alpha^{1/2}} \left(\frac{1}{2} + \alpha^{1/2}\eta + 2\alpha \tau \right) \right] e^{\alpha^{1/2}\eta} \\ &= \operatorname{rfe} \left[\frac{\eta}{2\tau^{1/2}} + (\alpha\tau)^{1/2} \right] + \left[\left(\frac{1 - a_{0}\Delta_{2}\eta + 2a_{0}^{2}\Delta_{2}^{2\tau}}{a_{0}^{2}\Delta_{2}^{2} - \alpha} - \frac{a_{0}\Delta_{2}}{(a_{0}^{2}\Delta_{2}^{-2} - \alpha)^{2}} \right) \chi_{1}a_{0} \\ &+ \frac{\chi_{2}a_{0}\Delta_{2}}{(a_{0}^{2}\Delta_{2}^{-2} - \alpha)} \right] \exp \left(-a_{0}\Delta_{2}\eta + (a_{0}^{2}\Delta_{2}^{-2} - \alpha) \tau \right) \operatorname{erfe} \left(\frac{\eta}{2\tau^{1/2}} - a_{0}\Delta_{2}\tau^{1/2} \right) \\ &+ \frac{\chi_{1}a_{0}}{(a_{0}^{2}\Delta_{2}^{-2} - \alpha)} \right] \exp \left(-a_{0}\Delta_{2}\eta + (a_{0}^{2}\Delta_{2}^{-2} - \alpha) \tau \right) \operatorname{erfe} \left(\frac{\eta}{2\tau^{1/2}} - a_{0}\Delta_{2}\tau^{1/2} \right) \\ &+ \frac{\chi_{1}a_{0}}{\pi^{1/2}} \left[\frac{2a_{0}\Delta_{2}\tau^{1/2}}{a_{0}^{2}\Delta_{2}^{-2} - \alpha} - \frac{1}{(a_{0}^{2}\Delta_{2}^{-2} - \alpha)^{2\tau^{1/2}}} + \frac{\chi_{3}\tau^{1/2}}{\chi_{1}a_{0}} \right] \exp \left(- \frac{\eta^{2}}{4\tau} - \alpha\tau \right) \right\}$$

$$(25.6)$$
For $\frac{b^{2}}{4} + \frac{a^{3}}{27} = 0$, and $\alpha = a_{0}^{2}\Delta_{2}^{-2}$, $\alpha \neq a_{0}^{2}\Delta_{1}^{-2}$
 $\theta(\eta, \tau) = \exp \left[-\frac{P_{\tau}}{2\tau} \left(a_{0} + \frac{P_{\tau}}{2a_{0}^{3}} \right) \eta - \frac{P_{\tau}}{2a_{0}^{4}} \left(e^{-a_{0}\eta} - \frac{1}{4} e^{-2a_{0}\eta} - \frac{3}{4} \right) \right] \cdot \left\{ \chi_{1}a_{0} \int_{0}^{\tau} \left[-2 \left(\frac{\alpha}{2\tau} \right)^{1/2} \left(a_{0}^{-1/2} - a_{0} \right)^{1/2} \right] \operatorname{exp} \left(- \frac{\eta^{2}}{4s} - \alpha \tau \right) \right\}$
 (25.6)
 $+ \left[\frac{\chi_{2}}{4(\alpha)^{1/2}} + \frac{\chi_{3}}{2(\alpha)^{1/2}} \left(a_{0}^{-1/2} - a_{0} \right) \eta - \frac{P_{\tau}}{2a_{0}^{4}} \left(e^{-a_{0}\eta} - \frac{1}{4} e^{-2a_{0}\eta} - \frac{3}{4} \right) \right] \cdot \left[\left(\chi_{1}a_{0} - \frac{1}{2} e^{-2a_{0}\eta} - \frac{3}{4} \right) \right] \right] \cdot \left[\left(\chi_{1}a_{0} - \frac{1}{2} e^{-2a_{0}\eta} - \frac{3}{4} \right) \right] ds$
 $+ \left[\frac{\chi_{2}}{4(\alpha)^{1/2}} + \frac{\chi_{3}}{2(\alpha)^{1/2}} \left(a_{0} \right) \right] \exp \left(- \alpha^{1/2}\eta \right) \right] ds$

18

· Sound from (25) by Intering S = 0. Dons, Set diana

$$+ x_{2} \left(\frac{\pi}{T} \right)^{1/2} \exp \left(- \frac{\pi^{2}}{4\pi} - \alpha \tau \right)$$

$$+ \frac{1}{2} \left(\frac{\pi^{2}}{2(\alpha)^{1/2}} \left(\frac{1}{2} + \alpha^{1/2} \eta + 2 \alpha \tau \right) + \frac{x_{3}}{2(\alpha^{1/2} + a_{0}A_{1})} \right] e^{-\alpha^{1/2} \eta} \operatorname{erfc} \left(\frac{\pi}{2 \tau^{1/2}} + (\alpha \tau)^{1/2} \right)$$

$$+ \frac{x_{3}^{a} \alpha_{A_{1}}}{a_{0}^{a} A_{1} - \alpha} \exp \left(-a_{0}A_{1} \eta + a_{0}^{2}A_{1}^{2} \tau - \alpha \tau \right) \operatorname{erfc} \left(\frac{\pi}{2\tau^{1/2}} - a_{0}A_{1} \tau^{1/2} \right) \right)$$

$$+ \frac{A_{2}^{2} + \frac{3}{2}A_{2} + \frac{1}{2} + \frac{P_{\tau}(1 - P_{\tau})}{4a_{0}} \left(1 + A_{2} \right) e^{-a_{0} \eta} + \frac{P_{\tau}(P_{\tau} - 2)}{16a_{0}} \left(A_{2} + \frac{1}{2} \right) e^{-2a_{0} \eta}$$

$$+ \frac{A_{2}^{2} - 2A_{1}A_{2} + \frac{3}{2} \left(A_{2} - A_{1} \right) - \frac{1}{2} + \frac{P_{\tau}(1 - P_{\tau})}{4a_{0}} \left(A_{2} - A_{1} - 1 \right) e^{-a_{0} \eta}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{16a_{0}} \left(A_{2} - A_{1} \right) e^{-2a_{0} \eta}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{16a_{0}} \left(A_{2} - A_{1} \right) e^{-2a_{0} \eta}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{16a_{0}} \left(A_{2} - A_{1} \right) e^{-2a_{0} \eta}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{16a_{0}} \left(A_{2} - A_{1} \right) e^{-2a_{0} \eta}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{16a_{0}} \left(A_{2} - A_{1} \right) e^{-2a_{0} \eta}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{16a_{0}} \left(A_{2} - A_{1} \right) e^{-2a_{0} \eta}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{16a_{0}} \left(A_{2} - A_{1} \right) e^{-2a_{0} \eta}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{16a_{0}} \left(A_{1} - A_{2} \right)^{2}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{(A_{1} - A_{2})^{2}} \left(A_{1} - A_{2} \right)^{2}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{(A_{1} - A_{2})^{2}} \left(A_{1} - A_{2} \right)^{2}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{(A_{1} - A_{2})^{2}} \left(A_{1} - A_{2} \right)^{2}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{(A_{1} - A_{2})^{2}} \left(A_{1} - A_{2} \right)^{2}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{(A_{1} - A_{2})^{2}} \left(A_{1} - A_{2} \right)^{2}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{(A_{1} - A_{2})^{2}} \left(A_{1} - A_{2} \right)^{2}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{(A_{1} - A_{2})^{2}} \left(A_{1} - A_{2} \right)^{2}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{(A_{1} - A_{2})^{2}} \left(A_{1} - A_{2} \right)^{2}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{(A_{1} - A_{2})^{2}} \left(A_{1} - A_{2} \right)^{2}$$

$$+ \frac{P_{\tau}(P_{\tau} - 2)}{(A_{1} - A_{2})^{2}} \left(A_{1} - A_{2} \right)^{2}$$

Eq. (25) represents the dimensionless transient temperature profile due to a step change in local heat flux. Since we are primarily interested in evaluating the wall temperature variation, the quantity desired is the dimensionless temperature at the wall $\theta(0,\tau)$ which can be readily formed from (25) by letting $\eta = 0$. Thus, for disks initially at adiabatic wall temperature, the wall temperature as a function dimensionless time is:

$$T_{w}(T) = T_{\infty} + \frac{\omega v}{c_{p}} \left[R^{2} S(0) + Q(0) \right] + \theta(0, T)$$
(26.a)

and for disks initially at uniform temperature, the same can be written:

$$T_{w}(\tau) = T_{iw} + \left(\frac{\nu}{w}\right)^{1/2} \frac{\Delta^{q}_{w}}{k} \theta(0,\tau)$$
(26.b)

The values of S(0) and Q(0) were obtained from the numerical integration of (12) and (13) by electronic computer, the Univac <u>1107</u>. Their value for Prandtl numbers ranging from 0.001 to 100 and suction parameter ranging from -1.2 to -4.0 were obtained.

Inasmuch (25) is valid for the entire time domain, the corresponding steady state solution at the wall can be obtained by simply letting $\tau \rightarrow \infty$, and $\eta \rightarrow 0$. The result is

$$\frac{b^2}{4} + \frac{a^3}{27} > 0$$
 and $\alpha \neq a_0^2 \beta_1^2$

- 1-

$$\theta_{ss}(0) = \frac{1}{a_0} \left[\frac{\Psi_1(0)}{\frac{\alpha^{1/2}}{a_0} - \beta_1} + 2 \operatorname{Re} \frac{\Psi_2(0) \left(\frac{\alpha^{1/2}}{a_0} - \beta_2 - i\beta_3 \right)}{\left(\frac{\alpha^{1/2}}{a_0} - \beta_2 \right)^2 + \beta_3^2} \right]$$

(27.a)

for
$$\frac{b^2}{4} + \frac{a^3}{27} > 0$$
 and $\alpha = a_0^2 \beta_1^2$ or $\alpha^{1/2} = -a_0^2 \beta_1$

$$\theta_{ss}(0) = \frac{1}{a_0} \left[\frac{\Psi_1(0)}{2\alpha^{1/2}} - 2 \operatorname{Re} \frac{\Psi_2(0) \left(\frac{-\alpha^{1/2}}{a_0} + \beta_2 + i\beta_3 \right)}{\left(\frac{\alpha^{1/2}}{a_0} - \beta_2 \right)^2 + \beta_3^2} \right]$$
(27.b)

for
$$\frac{b^2}{4} + \frac{a^3}{27} < 0$$
 and $\alpha \neq a_0^2 \Lambda_1^2$
 $\alpha \neq a_0^2 \Lambda_2^2$
 $\alpha \neq a_0 \Lambda_3^2$
 $\theta_{ss}(0) = \frac{\Phi_1(0)}{\alpha^{1/2} - a_0 \Lambda_1} + \frac{\Phi_2(0)}{\alpha^{1/2} - a_0 \Lambda_2} + \frac{\Phi_3(0)}{\alpha^{1/2} - a_0 \Lambda_3}$
(27.c)

for
$$\frac{b^2}{4} + \frac{a^3}{27} < 0$$
 and $\alpha = a_0^2 \Lambda_1^2$ (which automatically implies
 $\alpha \neq a_0^2 \Lambda_2^2$, $\alpha \neq a_0^2 \Lambda_3^2$)

$$\theta_{ss}(0) = \frac{\Phi_1(0)}{2\alpha^{1/2}} + \frac{\Phi_2(0)}{\alpha^{1/2} - a_0\Lambda_2} + \frac{\Phi_3(0)}{\alpha^{1/2} - a_0\Lambda_3}$$
(27.d)

For the cases $\alpha = a_0^2 \Lambda_2^2$ or $\alpha = a_0^2 \Lambda_3^2$ see footnote on page 15

for
$$\frac{b^2}{4} + \frac{a^3}{27} = 0$$
 and $\alpha \neq a_0^2 \Lambda_1^2$, $\alpha \neq a_0^2 \Lambda_2^2$

$$\theta_{ss}(0) = \frac{a_0 \chi_1(0)}{(\alpha^{1/2} - a_0 \Lambda_2)^2} + \frac{\chi_2(0)}{\alpha^{1/2} - a_0 \Lambda_2} + \frac{\chi_3(0)}{\alpha^{1/2} - a_0 \Lambda_1}$$
(27.e)

For
$$\frac{b^2}{4} + \frac{a^3}{27} = 0$$
 and $\alpha = a_0^2 \Lambda_1^2$

$$\theta_{ss}(0) = \frac{a_0 \chi_1(0)}{\left(\alpha^{1/2} - a_0 \Lambda_2\right)^2} + \frac{\chi_2(0)}{\alpha^{1/2} - a_0 \Lambda_2} + \frac{\chi_3(0)}{2 \alpha^{1/2}}$$
(27.f)

We have not worked the steady state case corresponding to Equation 25.g

In the above equations, the expressions $\Psi_n(0)$, Φ_n and $\chi_n(0)$ are obtained by letting $\eta = 0$ in the equations defining $\Psi(\eta)$ $\Phi_n(\eta)$ and $\chi_n(\eta)$.

For very large suction, i.e., $a_0 \rightarrow \infty$, the foregoing results simplify to:

$$\lim_{a_0 \to \infty} \theta_{ss}(0) = \frac{1}{a_0 P_r}$$
(28)

The numerical values of $\theta_{ss}(0)$ as evaluated from (27) are listed in Table 1 for the several Prandtl numbers and suction parameter shown. For comparison purposes, we have included the results for $P_r = 0.7$ obtained from the relation

$$\theta_{ss}(0) = -\frac{1}{\theta_{ss}'(0)}$$
 (Prime indicates first derivative with respect to η)

in which the values of $\theta_{ss}'(0)$ are taken from the results reported by Sparrow and Gregg [11]. The maximum discrepancy is about 2% for $a_0 = 1.2$ and the agreement is progressively better for large values of a_0 . It may also be shown that if one uses the limiting expression (28), the error remains within a few per cent when $a_0 > 2$.

22

Finally, we note that when dissipative effects are negligible, the steady wall temperature under the limiting condition of $a_0 \rightarrow \infty$ becomes

(29)

$$\lim_{m \to \infty} T_{w,ss} = T_{\infty} + \frac{q_w}{k} \left(\frac{v}{w} \right)^{1/2} \frac{1}{P_r^a_0} = T_{\infty} + \frac{q_w}{C_p \rho} |W_w|$$

a

The above results are to be expected on physical grounds because, when the suction is very large, the heat transfer at the wall would be completely dominated by the convective process. The effect of fluid injection $(a_0 < 0)$ is to decrease the heat transfer by blanketing the surface with the injecting fluid of the same temperature as T_w . On the other hand, suction increases heat transfer, because fluid at free stream temperature is effectively brought to the disk surface.

to the on a sample converter Principli instant tending this S, i et has an a time of the 4.0. This is illustrated in Y 5. 5. This tends are the for multicipathy large station the response time within the present power of the matter persenter. The same conclusion has also been drawn for the ones of a thry change in wall temperature [3]. Doe to the interity of the interity emploine, the contain families and be tending guartelized for any this thry wall hast flow which the Different's theorem.

4.0 Numerical Results and Discussion

The values of the roots of the cubic equation in the denominator of K are tabulated in Table III for various values of the Prandtl number and the blowing parameter.

The values of S(0) and Q(0) obtained from (12) and (13) are tabulated in Table II for the range of the Prandtl number and the blowing parameter as shown. It is seen from Table II that the Prandtl number and the blowing parameter has little effect on the values of S(0), which has almost a constant value of 0.5 in the range of P_r and positive a_0 under consideration. By contrast, Q(0) is a rapidly varying function of P_r and positive a_0 . It decreases with increasing P_r as well as a_0 . The ratio of the dimensionless temperature at the wall, $\theta(0,T) / \theta_{ss}(0)$, is shown plotted against $\tau = \frac{\omega t}{P_r}$ in Fig. 2 to 7 for the several Prandtl numbers indicated. If one replots the $\theta(0,\tau)/\theta_{ss}(0)$ against $a_0^2 P_r^2 (= a_0^2 P_r^{(0)})$ the data can, for all practical purposes, be brought to lie on a single curve for Prandtl number ranging from 0.1 to 100 and an from 1.5 to 4.0. This is illustrated in Fig. 8. This result indicates that for sufficiently large suction the response time varies inversely with the Prandtl number, the angular velocity and the second power of the suction parameter. The same conclusion has also been drawn for the case of a step change in wall temperature [9]. Due to the linearity of the energy equation, the previous results can be readily generalized for any arbitrary wall heat flux using the Duhamel's theorem.

5.0 A Special Case of Compressible Flow

For an ideal gas of constant Prandtl number, exhibiting a linear variation of viscosity with temperature, Ostrach and Thornton [16] has shown that, if dissipative effects can be ignored, the steady heat transfer solution for the compressible flow over a rotating disk can be obtained from the corresponding incompressible solutions. It is natural to examine if the same could be stated for the <u>unsteady</u> heat transfer processes. Using the same transformations for both the dependent and independent variables in the governing conservation equations as those given in [16], except for difference noted below, one may show that the answer to the equation posed above is affirmative. Now the velocity component normal to the disk is given by

$$w = \frac{\rho_{\infty}}{\rho} \left[(A \cup v_{\infty})^{1/2} H(\eta) - \frac{\partial}{\partial t} \int_{0}^{z} \frac{\rho}{\rho_{\infty}} dz \right]$$

Where the subscript ∞ refers to the free stream condition and A is the proportionality constant in the linear viscosity - temperature relation, i.e.

 $\frac{\mu}{\mu_{\infty}} = A \frac{T}{T_{\infty}}$

With the transformation stated above, the solution obtained for the incompressible case, may be directly applied for the compressible case. Expressions given for the transient wall temperature for the incompressible case remain valid provided that all properties are replaced by wall values.

REFERENCES

- Sparrow, E. M. and Gregg, J. L., "Non-Steady Surface Temperature Effects on Forced Convection Heat Transfer," <u>Journal of the Aeronautical Science</u>, 1957, vol. 24, pp. 776-777.
- Sparrow, E. M. and Gregg, J. L., "Prandtl Number Effects on Unsteady Forced-Convection Heat Transfer," NACA TN 4311, 1958.
- Cess, R. D., "Heat Transfer to Laminar Flow Across a Flat Plate with Nonsteady Surface Temperature," Journal of Heat Transfer, Trans. ASME, 1961, Series C, vol. 83, pp. 274-280.
- Riley, N., "Unsteady Heat Transfer for Flow Over a Flat Plate," Journal of Fluid Mechanics, Sept., 1963, vol. 17, Part I, pp. 97 - 104.
- Goodman, T. R., "Effect of Arbitrary Nonsteady Wall Temperature on Incompressible Heat Transfer," <u>Journal of Heat Transfer</u>, Trans. ASME, Series C, 1962, vol. 84, pp. 348 - 352.
- Adams, D. E. and Gebhart, B., "Transient Forced Convection From a Flat Plate Subjected to a Step Energy Input," <u>Journal of Heat Transfer</u>, Trans. ASME, Series C, May, 1964, vol. 86, pp. 253 - 259.
- Sparrow, E. M., "Unsteady Stagnation Point Heat Transfer," NASA TN, D-77 1595.
- Chao, B. T. and Jeng, D. R., "Unsteady Stagnation Point Heat Transfer," Journal of Heat Transfer, Trans. ASME, Series C, May, 1965, vol. 87, pp. 221 - 230.
- Jeng, D. R., "Unsteady Heat Transfer in Laminar Boundary Layers," Ph.D. Dissertation in Mechanical Engineering, University of Illinois at Urbana, 1965.
- Millsaps, K. and Pohlhausen, K., "Heat Transfer by Laminar Flow from a Rotating Plate," Journal of Aerospace Science, 1952, vol. 19, pp. 120 - 125.
- Sparrow, E. M. and Gregg, J. L., "Heat Transfer from a Rotating Disk to Fluids of any Prandtl Number," Trans. ASME, Series C, 1959, vol. 81, pp. 249 - 250.
- Sparrow, E. M. and Gregg, J. L., "Mass Transfer, Flow, and Heat Transfer about a Rotating Disk," Trans. ASME, Series C, 1960, vol. 82, pp. 294 - 302.
- Cess, R. D. and Sparrow, E. M., "Unsteady Heat Transfer from a Rotating Disk and at a Stagnation Point," International Developments in Heat Transfer, ASME, Part II, Section B, 1961, pp. 468 - 474.

- 14. Cochran, W. G., "The Flow due to a Rotating Disk," Proceedings of the Cambridge Philosophical Society, 1934, vol. 30, pp. 365 375.
- Stuart, J. T., "On the Effects of Uniform Suction on the Steady Flow due to a Rotating Disk," <u>Quarterly Journal of Mechanics and Applied</u> <u>Mathematics</u>, 1954, vol. 7, pp. 446 - 467.
- Ostrach, S. and Thornton, P. R., "Compressible Laminar Flow and Heat Transfer about a Rotating Isothermal Disk," NACA TN 4320, 1958.



Table 1

Values of the Dimensionless Temperature, $\theta_{\tt SS}(0)$, for Large Suction

| 100 | | 0,00747 | 0,00632 | 0.00487 | 0.00399 | 0.00334 | 0,00286 | 0.00250 | |
|------|----------------|---------|---------|---------|---------|---------|---------|---------|-------------------------|
| 10.0 | | 0.07745 | 0.07095 | 0.05015 | 0.03999 | 0.03333 | 0.02857 | 0.02500 | |
| | | | .0. | 50 | 1 | • | | 8 | 0 |
| 1.0 | | 0.7784 | 0.6466 | 0.4949 | 0.3983 | 0.3327 | 0.2854 | 0.2498 | |
| | From Ref. [11] | 1.1114 | 0.9200 | 0.7052 | 0.5685 | 0.4748 | | 0.3569 | |
| 0.7 | From Analysis | 1,0892 | 0.9151 | 0.7068 | 0.5685 | 0.4750 | 0.4076 | 0.3568 | 93 |
| 0.1 | | 7,0703 | 6,1733 | 4.8711 | 3.9566 | 3,3157 | 2.8490 | 2.4958 | 24 15 00 94x18 |
| 0.01 | | 68,6931 | 61,0106 | 48.5301 | 39,5055 | 33.1328 | 28.4784 | 24.9523 | 10 60 |
| Pr | a | 1.2 | 1.5 | 2.0 | 2.5 | 3.0 | 3.5 | 4.0 | |

Table 2 Values of S(0)

| P _r | | | a ₀ | | | |
|----------------|--------|--------|----------------|--------|--------|--------|
| | 1.2 | 1.5 | 2.0 | 2.5 | 3.0 | 4.0 |
| 0.01 | 0.4466 | 0.4657 | 0.4851 | 0.4939 | * | * |
| 0.1 | 0.4509 | 0.4720 | 0.4890 | 0.4952 | 0.4975 | 0.4992 |
| 0.7 | 0.4894 | 0.4942 | 0.4980 | 0.4992 | 0.4995 | 0.4998 |
| 1.00 | 0.4991 | 0.4998 | 0.5002 | 0.5001 | 0.5000 | 0.5000 |
| 10.0 | 0.5441 | 0.5228 | 0.5087 | 0.5038 | 0.5017 | 0.5005 |
| 100.0 | 0.5503 | 0.5258 | 0.5097 | 0.5042 | * | * |

*s failed to converge

| | a ₀ | | | | | | | |
|-------|------------------------|------------------------|------------------------|------------------------|-----------|------------------------|--|--|
| | 1.2 | 1.5 | 2.0 | 2.5 | 3.0 | 4.0 | | |
| 0.01 | 9633 | 7106 | 4560 | 3093 | * | * | | |
| 0.1 | 96.9393 | 75.1505 | 48.3896 | 32.4440 | 22.9324 | 13.0515 | | |
| 0.7 | 2.9213 | 2.1302 | 1.3088 | 0.8624 | 0.6055 | 0.3431 | | |
| 1.0 | 1.6548 | 1.1874 | 0.7199 | 0.4718 | 0.3307 | 0.1872 | | |
| 10.0 | 0.1159 | 0.0660 | 0.0329 | 0.0200 | 0.1362x10 | 10.756×10^{2} | | |
| 100.0 | 0.4598x10 ¹ | 0.1924x10 ¹ | 0.5911x10 ² | 0.2594x10 ² | * | * | | |

Values of Q(0)

* s failed to converge

Table 3 Roots of the Cubic Equation:

.5000 .3978 - 1.0000 6665. - 1.0005 .4988 .4990 .5000 .0050 .0501 3507 - 5.0096 - 1,0001 - 1.0007 - 1.0242 7.0155 -50.0851 - 1.1547 4.0 . 1 1 1 . 1 1 1 1 .5000 .0502 .0050 .4998 .3512 .4980 .4983 - 1.0012 .3300 - 1.0000 - 1.0008 .5000 - 5.0162 -50.1330 12.2049 - 1.1601 - 1.0001 - 1.0371 3.5 1 . 1 1 1 1 1 1 1 .3522 6967. .5000 .5000 7664. .4963 - .1974 - 1.1638 - 1.0002 - 1.0015 - 5.0293 .0050 - 1.0000 - 1.0023 - 1.0572 22.8684 -50.2104 .0503 3.0 . 1 1 . 1 1 1 1 .4922 .0924 -50.3256 .0506 .4937 .5000 - 5.0578 - 1.0866 - 1.1660 - 1.0000 6665* .4993 .3547 - 1.0047 47.6720 .0051 - 1.0003 - 1.0030 2.5 . (Blowing Parameter) 1 . 1 1 1 1 1.0072 .5000 .8414 .4998 .0516 4983 .3625 4800 .4847 1.0114 5.1242 1.1234 116.5402 - 50.4663 - 1.1670 .0052 1.0008 1.0001 2.0 1 1 1 1 1 1 1 1 1 1 1 1 1 ao .5000 50.5876 .4945 .4136 0312 .4538 1.1578 368,1564 - 1.1662 4995 1.0339 3.3904 .0055 1.0002 .0550 1.0026 1.0220 5.2757 1.5* . 1 1 1 1 ı . ı 1 1 1 1 1 . .5000 .4865 8.5302 - 1.1634 1.0006 898.4965 - 50.6339 5.4190 .0062 .4988 .0623 .3990 7960 .3951 1.0747 1.1717 1.0062 1.0500 1.2* 1 1 1 1 1 1 1 1 1 1 1 1 1 1 Root 4ª Y A2 R V A2 A2 A3 V A2 R T A T A2 r 4 T 10.00 1.00 100.00 0.10 0.70 10.0 PR

*For Prandtl number 0.7 and blowing parameters 1.2 and 1.5, the names $^{\Lambda}$ 1, $^{\Lambda}$ 2, and $^{\Lambda}$ 3 are to be replaced by β_1 , β_2 , and β_3 respectively.





TO A STEP CHANGE IN LOCAL HEAT FLUX







TO A STEP CHANGE IN LOCAL HEAT FLUX



TO A STEP CHANGE IN LOCAL HEAT FLUX



. ,



50 cepies received by Madanglais 5 Dr. I.T. K. Smull 1 Dr. Herman 1 Mrs. Mª Caulees Mr. John M- Daniel 1 Sorgas Library 1 Loy A. Bland John Mc Clurkin Office of Shanter V heclarch Centracto Mies Barbara Danis Dr. N.H. Kurzweg

